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High-dimensional neural network potentials for magnetic systems using spin-dependent atom-centered symmetry functions

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Machine learning potentials have emerged as a powerful tool to extend the time and length scales of first-principles quality simulations. Still, most machine learning potentials cannot distinguish different electronic spin arrangements and thus are not applicable to materials in different magnetic states. Here we propose spin-dependent atom-centered symmetry functions as a type of descriptor taking the atomic spin degrees of freedom into account. When used as an input for a high-dimensional neural network potential (HDNNP), accurate potential energy surfaces of multicomponent systems can be constructed, describing multiple collinear magnetic states. We demonstrate the performance of these magnetic HDNNPs for the case of manganese oxide, MnO. The method predicts the magnetically distorted rhombohedral structure in excellent agreement with density functional theory and experiment. Its efficiency allows to determine the Néel temperature considering structural fluctuations, entropic effects, and defects. The method is general and is expected to be useful also for other types of systems such as oligonuclear transition metal complexes.

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INTRODUCTION

In recent years, machine learning potentials (MLPs), which allow extending the time and length scales of first-principles quality atomistic simulations^{1–3}, have become a rapidly growing field of research. Increasingly complex systems have been investigated, driving developments and extending the applicability of MLPs. Many of the current MLPs can be classified into four generations as follows^{4,5}: the first generation of MLPs proposed already in 1995⁶ typically employs a single or a few feed-forward neural networks, making them applicable to low-dimensional systems^{7,8}. In 2007, second-generation MLPs have become available with the introduction of high-dimensional neural network potentials (HDNNP)^{9–12}. These MLPs are applicable to systems containing thousands of atoms by making use of the locality of a large part of the atomic interactions. For this purpose, the potential energy is calculated as a sum of local environment-dependent atomic energies defined by a cutoff radius. Most modern MLPs belong to this second generation, e.g., HDNNPs^{9,13}, Gaussian approximation potentials¹⁴, moment tensor potentials (MTPs)¹⁵, and many others^{16–18}. The third generation of MLPs includes long-range interactions, mainly electrostatics but also dispersion, beyond the cutoff radius. The electrostatic interactions can be based on element-specific fixed charges^{14,19} or local environment-dependent atomic charges expressed by machine learning (ML)^{20–22}. Finally, fourth-generation MLPs take non-local or even global dependencies in the electronic structure into account, and consequently the atomic charges can adapt to non-local charge transfer and even different global charge states. The first method of this generation has been the charge equilibration neural network technique²³ and further methods such as Becke population neural networks (BpopNN)²⁴ and fourth-generation HDNNPs²⁵ have been introduced recently.

In spite of these advances, which extended the complexity of the systems and the physical phenomena that can be studied, a remaining limitation of MLPs is the inability to take different spin arrangements and thus magnetic interactions into account. The reason for this limitation is that typically MLPs are trained to represent the potential energy surface (PES) of one electronic state as a function of structural descriptors—most often but not necessarily the ground state of a system. With the exception of different global charge states in fourth-generation potentials, describing multiple electronic states usually requires to train separate MLPs for each state^{26–30} or to use one MLP yielding a set of output energies corresponding to the different excited states^{31–34}. The unique structure–energy relation of a given state is a central component of nowadays MLPs and energy changes resulting from different spin arrangements give rise to contradictory information in the training process. Only very recently, magnetic MTPs (mMTPs)³⁵ have been proposed as an example of an MLP, which is able to describe magnetic systems containing a single element such as iron.

For studying magnetic materials, the description of a single magnetic state is insufficient, as the atomic magnetic moments fluctuate at finite temperatures. Atomic spin flips often play a role already at ambient temperatures, as the energy differences between different spin configurations are typically in the order of a few meV atom^{−1} only. For example, the Curie and Néel temperatures, at which a material loses its ferromagnetic or antiferromagnetic state, are typically below 1000 and 500 K, respectively³⁶. Magnetic transitions often give rise to structural changes³⁷. However, current implementations of MLPs are not able to capture these effects, because the magnetic interaction is not included explicitly. For example, if an antiferromagnetic ground state is used for training the MLP, this state will be retained in simulations at higher temperatures irrespective of the magnetic ground state under these conditions.

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To calculate the energy of a magnetic configuration in an atomistic simulation, models such as the Ising model³⁸, the Heisenberg model³⁹, and the Hubbard model⁴⁰ are widely used. However, these models are based on lattices, and structural and spin changes at finite temperatures cannot be considered simultaneously. Conventional interatomic potentials, which are able to take into account both contributions, have been developed only for a few systems such as iron^{41–43}. The only generally applicable option is the use of ab initio molecular dynamics, which is based on the explicit calculation of the electronic structure in each step and thus is able to distinguish different magnetic states. This approach, however, is inevitably associated with high computational costs restricting simulations to small systems and short timescales.

Beyond the field of PESs, several ML approaches have been developed to address the properties of magnetic compounds, e.g., for the prediction of magnetic moments³⁶ and ordering temperatures^{44–46} from structural parameters. Moreover, ML methods are able to classify ferromagnetic and antiferromagnetic ground-state materials⁴⁷ and to predict spin-state splittings and metal–ligand bond distances in transition metal complexes^{48,49}. Recently, a high-dimensional neural network approach for the prediction of oxidation and spin states (high-dimensional neural network spin prediction (HDNNS)) has been developed⁵⁰ and also the Atoms-in-Molecules Network allows predicting atomic spins⁵¹.

In spite of these applications of ML to magnetic compounds, with the exception of mMTPs³⁵, to date there is no method enabling large-scale first-principles quality atomistic simulations of systems explicitly including magnetic interactions and different magnetic states. The main reason is that current MLPs use descriptors exclusively depending on the atomic structure, such as atom-centered symmetry functions (ACSF)⁵², smooth overlap of atomic positions⁵³, and many others⁵⁴. Exceptions are BpopNNs²⁴ and the recently introduced fourth generation of HDNNS²⁵ in which the atomic charges are included as additional information besides the structural descriptors, to provide qualitative information about the electronic structure.

Describing the spin configuration in a form suitable as input for MLPs is very challenging, as not only the absolute spin values but also the relative spin-up and spin-down arrangements and the relative positions of the atoms are vital. Hence, to predict the energy and forces as a function of the geometric and spin configuration, spin-dependent descriptors are needed. Here we propose spin-dependent ACSFs (sACSFs) to construct MLPs simultaneously applicable to multiple magnetic states including excited states. The sACSFs add the description of the magnetic configuration for the case of collinear spin polarization corresponding to spin-polarized density functional theory (DFT) calculations, i.e., considering spin-up and spin-down electrons without an associated spatial direction. They produce atomic energies as a function of the local geometric and spin environment, and thus formally represent a second-generation potential that can also be combined with third- and fourth-generation MLPs to include additional physics such as long-range electrostatic interactions. In this work, we will benchmark sACSFs employing second-generation HDNNS⁹, extending these potentials by an explicit dependence on the full collinear spin configuration space to describe magnetic interactions.

We choose manganese oxide, MnO, to assess the quality of the resulting magnetic HDNNS (mHDNNS) that can be constructed using sACSFs, because of the well-characterized antiferromagnetic ground-state configuration^{55,56}, the Néel temperature of $T_N^{\text{exp}} = 116 \text{ K}$ ^{57,58}, and the rhombohedral distortion of the antiferromagnetic phase with lattice constant $a^{\text{exp}} = 4.430 \text{ \AA}$ and lattice angle $\alpha^{\text{exp}} = 90.62^\circ$ at 8 K ⁵⁹, making this system a very interesting and challenging benchmark case. The magnetic unit cell is a $2 \times 2 \times 2$ supercell of the geometric unit cell and is built from (111) planes of ferromagnetically coupled Mn ions⁵⁶.

These planes couple antiferromagnetically to the neighboring planes. Spin-polarized DFT calculations employing the hybrid functional PBE0^{60,61} yield the correct magnetic ground state, called antiferromagnetic type II (AFM-II), with the rhombohedral lattice parameters $a^{\text{PBE0}} = 4.40 \text{ \AA}$ and $\alpha^{\text{PBE0}} = 90.88^\circ$ in good agreement with the experiment⁶². The rhombohedral distortion is a consequence of the magnetic anisotropy, which arises from the magnetic dipole interactions⁶³.

Using the mHDNNS, we are able to perform first-principles quality high-throughput studies of the equilibrium geometries and energetics of different collinear magnetic configurations. We can simulate the antiferromagnetic to paramagnetic phase transition to determine the Néel temperature and the associated structural changes. Moreover, the mHDNNS enables the inclusion of defects, and in simulations including Mn vacancies we investigate the formation of different oxidation states by using an additional HDNNS to analyze the temperature dependence of the magnetic moment of the obtained ferrimagnetic configuration.

RESULTS AND DISCUSSION

Magnetic high-dimensional neural network potential

The reference data set of the mHDNNS consists of 3101 $2 \times 2 \times 2$ bulk supercells of MnO and $\text{Mn}_{0.969}\text{O}$ in various magnetic states and their corresponding HSE06^{64–66} DFT energies and forces. The supercells include different displacements of the atomic positions from the ideal rock salt lattice and distortions of the lattice parameters as well. The construction of the reference data set is described in detail in the Supplementary Methods^{67–69}.

Training this data set with conventional ACSFs yields an energy root mean squared error (RMSE) of about 11 meV atom^{-1} in the best potential we were able to construct. This RMSE is an order of magnitude larger than the usual HDNNS accuracy of 1 meV atom^{-1} , because the ACSFs can only provide geometrical information to assign the energy. Thus, HDNNS based on ACSFs only predict an averaged potential energy ignoring the magnetic configurations, whereas e.g., the HSE06 DFT functional yields an energy difference of $45.9 \text{ meV atom}^{-1}$ between the AFM-II order and the ferromagnetic (FM) order for the ideal rock salt MnO structure using the experimental lattice constant a^{exp} .

By including sACSFs to distinguish the magnetic configurations, the energy RMSE is reduced by one order of magnitude to about 1 meV atom^{-1} . The energy difference between the AFM-II and FM orders is resolved and is predicted to be $46.3 \text{ meV atom}^{-1}$ in excellent agreement with the HSE06 reference. Both results demonstrate the ability of the sACSFs to accurately describe the different magnetic configurations. The energy and force component prediction errors are plotted as a function of the reference values in Fig. 1a, b. The underlying number of Mn_xO reference structures and key performance indicators such as RMSEs, maximum absolute errors, and fractions of data points with high errors are compiled in Table 1 for the training and test data sets. Even when including various magnetic states, the accuracy of the mHDNNS is in the typical region of state-of-the-art MLPs of 1 meV atom^{-1} and $0.1 \text{ eV } a_0^{-1}$ for energies and forces with a test set energy RMSE of $1.11 \text{ meV atom}^{-1}$ and a test set force components RMSE of $0.066 \text{ eV } a_0^{-1}$. a_0 is the Bohr radius. In comparison, the RMSE of a recent mMTP for defect-free body centered cubic iron restricted to a fixed lattice parameter is $2.0 \text{ meV atom}^{-1}$ ³⁵, i.e., in the same order of magnitude.

The absolute atomic spin values $|M_S|$, equivalent to the half-difference of the number of spin-up and spin-down electrons, can be predicted by a separate HDNNS. This HDNNS is trained on the same reference structures covering the range of $1.7 < |M_S| < 2.4$ in case of Mn spins. The resulting RMSE value for these Mn spins is 0.03 for both the training and test set. Only 0.25% of the test data exhibits an error > 0.1 , whereby the maximum absolute error is

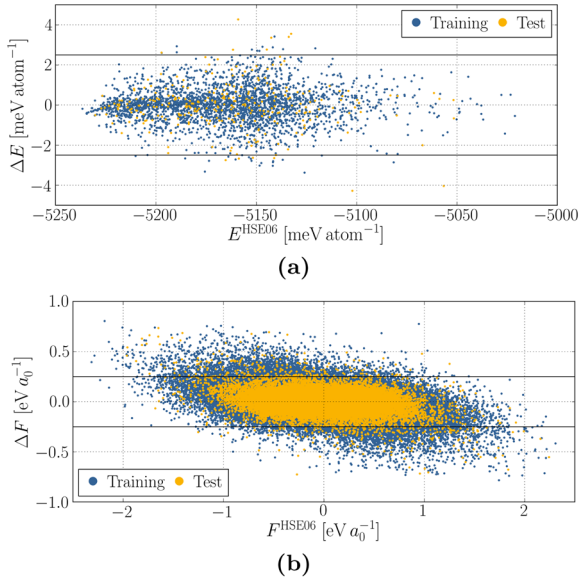


Fig. 1 Accuracy of the mHDNNP. a Energy errors $\Delta E = E^{\text{mHDNNP}} - E^{\text{HSE06}}$ as a function of the reference energy E^{HSE06} and **b** force component errors $\Delta F = F^{\text{mHDNNP}} - F^{\text{HSE06}}$ as a function of the reference force components F^{HSE06} .

Table 1. Details of the mHDNNP fit.		
	Training set	Test set
$N_{\text{struct}}(\text{MnO})$	1387	156
$N_{\text{struct}}(\text{Mn}_{0.969}\text{O})$	1421	137
RMSE(E)	0.86	1.11
ΔE_{max}	3.4	4.3
$\Delta E > 2.5$	1.07%	4.44%
RMSE(F)	0.067	0.066
ΔF_{max}	0.81	0.73
$\Delta F > 0.25$	0.98%	0.89%

Number of Mn_xO structures N_{struct} in the training and test set, and the key performance indicators RMSE, maximum absolute error, and fraction of data points with energy and force errors of the mHDNNP higher than 2.5 meV atom⁻¹ and 0.25 eV a₀⁻¹, respectively. The reference energy range is $-5236.6 \text{ meV atom}^{-1} \leq E \leq -5021.2 \text{ meV atom}^{-1}$ and the reference force components range is $|F| \leq 2.31 \text{ eV a}_0^{-1}$. Energies are provided in meV atom⁻¹ and force components in eV a₀⁻¹.

0.20 (Supplementary Fig. 1). Consequently, information about different Mn oxidation states and the magnetic moments can be obtained with accuracy comparable to the underlying first-principles method.

Magnetic configurations

Besides the accurate description of the energetics, the sACSFs enable to predict structural changes arising from the magnetic configuration. For example, for the rhombohedral MnO global minimum structure with AFM-II order (Fig. 2a), an unconstrained optimization yields the lattice parameters $a^{\text{mHDNNP}} = 4.433 \text{ \AA}$ and $\alpha^{\text{mHDNNP}} = 90.77^\circ$ in very good agreement with the HSE06 DFT results $a^{\text{HSE06}} = 4.434 \text{ \AA}$ and $\alpha^{\text{HSE06}} = 90.89^\circ$, as well as experimental data $a^{\text{exp}} = 4.430 \text{ \AA}$ and $\alpha^{\text{exp}} = 90.62^\circ$ ⁵⁹.

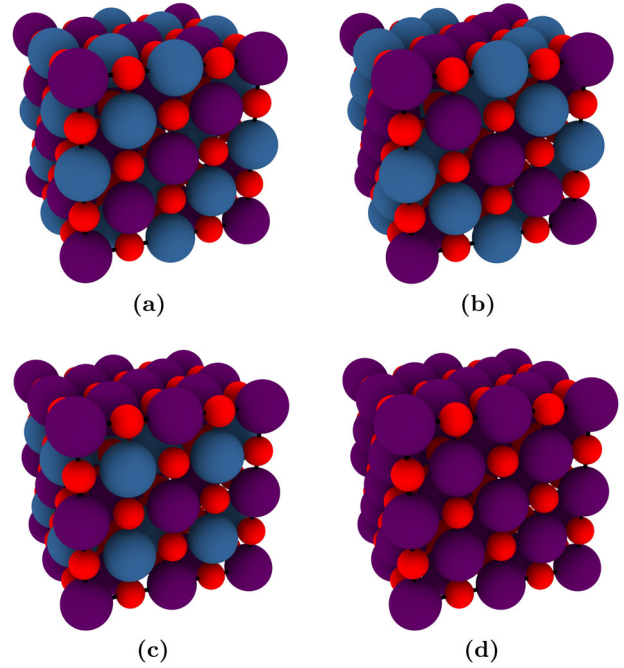


Fig. 2 Different magnetic orders of MnO. a Rhombohedral MnO global minimum structure with AFM-II order, **b** magnetic order degenerate to the cubic AFM-II order, **c** AFM-I order, and **d** FM order of MnO. Small red balls represent O atoms, large violet ones Mn atoms with spin-up, and large blue ones Mn atoms with spin-down⁹⁵. Different perspectives of the magnetic orders **a** and **b** are shown in the Supplementary Fig. 3.

Further, also excited magnetic configurations are predicted in excellent agreement with HSE06. For example, the lattice parameter of the resulting cubic lattice of the FM configuration (Fig. 2d) is $a_{\text{FM}}^{\text{mHDNNP}} = 4.461 \text{ \AA}$ compared to $a_{\text{FM}}^{\text{HSE06}} = 4.462 \text{ \AA}$. The energy difference to the global minimum is 45.8 meV atom⁻¹ in excellent agreement with the HSE06 value 45.5 meV atom⁻¹. The longer distances between ferromagnetically compared to antiferromagnetically interacting Mn ions in MnO are emphasized by the optimized lattice parameters of the AFM-I configuration (Fig. 2c) with tetragonal lattice parameters $a_{\text{AFM-I}}^{\text{mHDNNP}} = 4.461 \text{ \AA}$ and $c_{\text{AFM-I}}^{\text{mHDNNP}} = 4.414 \text{ \AA}$ ($a_{\text{AFM-I}}^{\text{HSE06}} = 4.459 \text{ \AA}$ and $c_{\text{AFM-I}}^{\text{HSE06}} = 4.420 \text{ \AA}$), as the lattice is elongated in both directions of FM interactions to the nearest neighbors. In conclusion, the lattices of FM (cubic), AFM-I (tetragonal), and AFM-II orders (rhombohedral), as well as other magnetic configurations can be distinguished due to the magnetic interaction. Further, the mHDNNP can provide information about the influence of defects on the magnetic order. For example, one spin-flip in the AFM-II configuration per $2 \times 2 \times 2$ supercell reduces the rhombohedral distortion by 0.09° (HSE06: 0.11°). The corresponding energy increase is 1.5 meV atom⁻¹ (HSE06: 1.5 meV atom⁻¹).

The efficiency of the mHDNNP in combination with a basin-hopping Monte Carlo (MC) scheme⁷⁰ in which MC spin flips are employed instead of atomic displacements enables high-throughput searches of the minima in spin configuration space, which would be computationally too demanding employing DFT. The spin-flip basin-hopping MC simulations of $2 \times 2 \times 2$ MnO supercells and molecular dynamics (MD) simulations including MC spin flips (MDMC) of $6 \times 6 \times 6$ MnO supercells confirmed the rhombohedral AFM-II magnetic order to be the global minimum in agreement with experiments⁵⁶. However, if the lattice is restricted to be cubic, two degenerate global minima exist: the AFM-II configuration shown in Fig. 2a and another antiferromagnetic

order shown in Fig. 2b, which is not based on (111) planes of ferromagnetically coupled Mn ions. DFT calculations confirm this result. Radial distribution functions (Supplementary Fig. 4) show that the same ferromagnetic and antiferromagnetic interactions are present in the two configurations. This second cubic global minimum configuration stays cubic in an unconstrained optimization. The rhombohedral distortion of the AFM-II order can therefore be identified as the origin of the energetic preference. The energy gain by the rhombohedral distortion is $1.9 \text{ meV atom}^{-1}$ (HSE06: $2.1 \text{ meV atom}^{-1}$).

Magnetic interactions

For the ideal cubic structure, a Heisenberg spin Hamiltonian can be constructed,

$$H = -\frac{J_1 k_B}{2} \sum_i \sum_{n_i} \mathbf{S}_i \mathbf{S}_{n_i} - \frac{J_2 k_B}{2} \sum_i \sum_{m_i} \mathbf{S}_i \mathbf{S}_{m_i}, \quad (1)$$

which includes the magnetic interactions of atom i with its nearest neighbors n_i and next nearest neighbors m_i . The strengths of the magnetic coupling between the vector spin operators \mathbf{S}_i and \mathbf{S}_{n_i} , as well as \mathbf{S}_i and \mathbf{S}_{m_i} , are given by the exchange coupling constants J_1 and J_2 , respectively. k_B is the Boltzmann constant. The exchange coupling constants can be determined employing energy differences (per atom) between the FM, AFM-I, and AFM-II configurations. Rearranging Eq. (1) for these systems yields

$$J_1 = \frac{E_{\text{AFM-I}} - E_{\text{FM}}}{4S^2 k_B}, \quad (2)$$

$$J_2 = \frac{4E_{\text{AFM-II}} - 3E_{\text{AFM-I}} - E_{\text{FM}}}{12S^2 k_B}, \quad (3)$$

with the spin $S = \frac{5}{2}$ of the high-spin Mn^{II} ions (Supplementary Discussion). Using mean field theory⁷¹, the Néel temperature can be calculated as

$$T_N = -2S(S+1)J_2. \quad (4)$$

Employing the HSE06 DFT energies for the lattice constant a^{exp} , we obtain $J_1^{\text{HSE06}} = -13.9 \text{ K}$, $J_2^{\text{HSE06}} = -14.5 \text{ K}$, and $T_N^{\text{HSE06}} = 255 \text{ K}$. The mHDNNP results match these values almost perfectly with $J_1^{\text{mHDNNP}} = -14.0 \text{ K}$, $J_2^{\text{mHDNNP}} = -14.6 \text{ K}$, and $T_N^{\text{mHDNNP}} = 256 \text{ K}$. This agreement underlines the accuracy of the sACSFs and of the mHDNNP method in describing the multiple PESs of the MnO magnetic states.

The overestimation of HSE06 compared to the experimental Néel temperature $T_N^{\text{exp}} = 116 \text{ K}$ ^{57,58} was also found for other hybrid DFT functionals in previous studies. For example, the PBE0 functional yields $T_N = 240 \text{ K}$ ⁶² and the HSE03 functional $T_N = 230 \text{ K}$ ⁷². Similar to these functionals, HSE06 predicts too negative exchange coupling constants of MnO compared to the experimentally determined values $J_1^{\text{exp}} = -8.5 \text{ K}$ and $J_2^{\text{exp}} = -9.6 \text{ K}$ ^{73,74}. Therefore, the strength of antiferromagnetic interactions is overestimated favoring the stability of the antiferromagnetic phase. The Néel temperature is consequently overestimated by the mHDNNP as well, which is based on HSE06 reference data. This overestimation is thus not an inherent error of the mHDNNP method and the agreement with experiment could be improved by using another reference method.

Néel temperature

Mean field theory misses the influence of specific magnetic configurations of MnO on the Néel temperature. Moreover, the underlying Heisenberg spin Hamiltonian restricts the system to a fixed lattice. Employing the mHDNNP, we can overcome both limitations step by step to reveal their influences on the Néel temperature. Conventional MC spin-flip simulations use a fixed lattice but allow exploring the specific states of MnO. Using

isothermal-isobaric (NpT) MD trajectories in MDMC simulations, both the atomic positions and the lattice parameters can adapt to the magnetic configurations including thermal fluctuations.

The phase transition temperature is observable as peak in the molar heat capacity as a function of the temperature. The molar heat capacity at constant volume C_V can be obtained from MC spin-flip simulations and the molar heat capacity at constant pressure C_p from MDMC simulations as described in section “Simulation analysis.”

To identify the AFM-II to PM transition, the temperature dependence of the order parameter C defined in section “Simulation analysis” can be used. This order parameter takes into account the maximal alignment of the magnetic configurations to an AFM-II configuration. The order parameter is $C = 1$ if the magnetic configuration corresponds to AFM-II during the entire simulation. It is $C = 0.5$ in case of the second cubic global minimum. The variety of paramagnetic (PM) configurations leads to a smaller value $C \lesssim 0.2$, as the PM orders are not correlated to the AFM-II order.

MC spin-flip simulations of a cubic $6 \times 6 \times 6$ MnO supercell using the lattice constant a^{exp} yield a transition temperature of 300 K ($\text{MC}_{\text{cub}}^{\text{AFM-II}}$ and MC_{cub} in Fig. 3). We note that these simulations are different from MC simulations using only J_1 and J_2 coupled interactions, because here the full PES from the HSE06 reference data is explored. The order parameter of these simulations ($\text{MC}_{\text{cub}}^{\text{AFM-II}}$ and MC_{cub} in Fig. 4) proves that the transition temperature belongs to the antiferromagnetic to PM transition. The Néel temperature is the same for the disordering process ($\text{MC}_{\text{cub}}^{\text{AFM-II}}$) and the ordering process (MC_{cub}), i.e., no hysteresis is present. The AFM-II to PM ($\text{MC}_{\text{cub}}^{\text{AFM-II}}$) and the second cubic global minimum to PM ($\text{MC}_{\text{cub}}^{\text{nd}}$) transition temperatures are identical because of the equal interactions of both cubic global minimum configurations. The ordering process can consequently also end in the second cubic global minimum configuration as had happened in the MC_{cub} simulations at 120 , 160 , and 200 K . However, below the Néel temperature, the interconversion of both cubic global minima is kinetically hindered. In summary, the inclusion of specific state information increases the Néel temperature of MnO by about 50 K compared to the mean field theory.

The restriction to a cubic lattice is an approximation, because the AFM-II minimum energy configuration has a rhombohedral

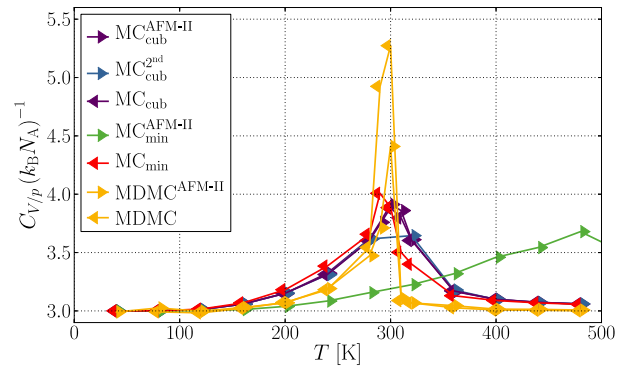


Fig. 3 Temperature dependence of the heat capacity. Molar heat capacity at constant volume C_V (MC) and pressure C_p (MDMC) divided by the Boltzmann constant k_B and the Avogadro constant N_A as a function of the temperature T obtained in different simulation methods for $6 \times 6 \times 6$ MnO supercells. The superscript of the simulation method defines the initial magnetic configuration (cubic AFM-II or second global minimum), otherwise random initial magnetic orders are used. The subscript cub indicates the restriction to the cubic lattice with $a^{\text{exp}} = 4.430 \text{ \AA}$ and the subscript min indicates an optimization of the initial structure prior to the MC simulation.

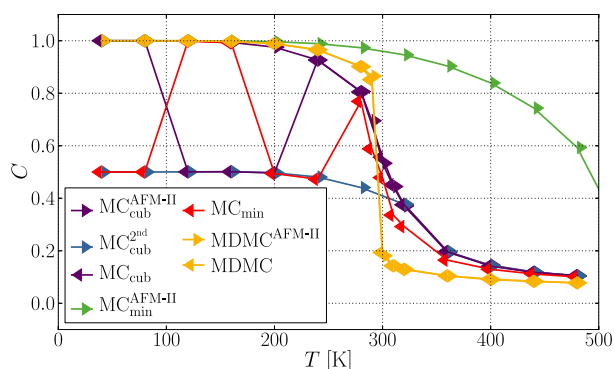


Fig. 4 Temperature dependence of the order parameter. Order parameter C as a function of the temperature T obtained in different simulation methods for $6 \times 6 \times 6$ MnO supercells. The superscript of the simulation method defines the initial magnetic configuration (cubic AFM-II or second global minimum), otherwise random initial magnetic orders are used. The subscript cub indicates the restriction to the cubic lattice with $a^{\text{exp}} = 4.430 \text{ \AA}$ and the subscript min indicates an optimization of the initial structure prior to the MC simulation. Lines are only drawn to guide the eyes.

lattice. Employing the rhombohedral lattice ($\text{MC}_{\text{min}}^{\text{AFM-II}}$) increases the Néel temperature to about 480 K. This increase can be explained by the mean energies of the AFM-II and PM configurations. Although the energy difference is $7.0 \text{ meV atom}^{-1}$ for the experimental cubic lattice, it is $10.6 \text{ meV atom}^{-1}$ for the rhombohedral lattice. To obtain similar Boltzmann factors during the $\text{MC}_{\text{min}}^{\text{AFM-II}}$ simulation compared to the $\text{MC}_{\text{cub}}^{\text{AFM-II}}$ simulation, the temperature has to be raised by the same factor of 1.5, which is similar to the increase of the Néel temperature (Supplementary Discussion). The data of the PM phase were calculated from 1000 PM configurations, whose magnetic orders were obtained during the 10 ns MDMC simulation at 400 K. On the other hand, an optimized lattice of the initial PM configuration (MC_{min}) leads to a Néel temperature of about 290 K. In conclusion, the choice of a specific fixed lattice can change the Néel temperature by about 200 K. Adapting the lattice and the atomic positions to the magnetic order is consequently of major importance to obtain reliable results.

NpT MD enables to sample the thermodynamic equilibrium of the given magnetic configuration at pressure p and temperature T to get rid of simulation artifacts due to a restricted geometric structure. The MDMC simulations are therefore a much better representation of experimental conditions. MDMC simulations predict an AFM-II to PM transition at 300 K (MDMC^{AFM-II} and MDMC in Figs. 3 and 4). This temperature is similar to the result of the cubic lattice due to a compensation of two main factors: the optimized AFM-II lattice (at $p = 1$ bar) leads to an energy gain of $1.9 \text{ meV atom}^{-1}$ compared to the cubic lattice with a^{exp} , whereas the mean energy gain of the PM configurations is $1.3 \text{ meV atom}^{-1}$. In addition, thermal fluctuations of the atomic positions are included in MDMC simulations leading to thermal expansion of the MnO lattice with increasing temperature. To quantify the energy increase due to the larger lattice volume, optimizations at various pressures of the AFM-II and PM configurations were performed. If the optimized volume at $p = 1$ bar is increased to the volume given at 300 K in the MDMC simulations (Fig. 5), the energy of the AFM-II configuration is increased by $1.7 \text{ meV atom}^{-1}$ and the mean energy of the PM configurations by $1.1 \text{ meV atom}^{-1}$ (Supplementary Fig. 5^{75,76}).

The cube root of the mean lattice volume, i.e., a hypothetical averaged cubic lattice constant, shows a discontinuous increase of 0.005 \AA at the Néel temperature (Fig. 5). This increase is similar to the difference of the optimized AFM-II and PM values, which also differ by 0.005 \AA . In accordance with the simulations,

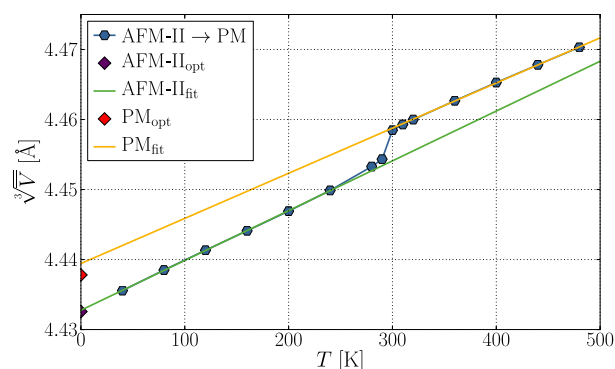


Fig. 5 Thermal expansion of MnO. Cube root of the mean volume per unit cell $\sqrt[3]{V}$ as a function of the temperature T obtained in 10 ns NpT MD simulations including MC spin flips of $6 \times 6 \times 6$ MnO supercells and the optimized values at $p = 1$ bar of the AFM-II and PM configurations. The straight lines represent linear fits for the high- and low-temperature regions.

experimentally an increase of about 0.004 \AA has been observed at the Néel temperature⁷⁷. The mean lattice angle decreases from 90.77 to 90.00° at the Néel temperature as shown in the Supplementary Fig. 6 matching the experimental PM angle of 90.00° ⁷⁷. This discontinuous change of the lattice volume and shape confirms a first-order magnetic phase transition^{78,79}.

The linear thermal expansion coefficient of the PM phase has been measured to be $\alpha_L^{\text{exp}} = 12 \cdot 10^{-6} \text{ K}^{-1}$ at 400 K⁸⁰. Employing the data from the MDMC simulations of the PM phase $\alpha_L^{\text{mHDNNP}} = 14 \cdot 10^{-6} \text{ K}^{-1}$ at 400 K is obtained in good agreement with the experiment.

In conclusion, mHDNNPs combine the accuracy and generality of first-principles methods with an efficiency close to spin lattice models. The Néel temperature of MnO calculated by mean field theory differs by about 50 K from the MC result, which explicitly samples the magnetic configurations. Including structural fluctuations in the prediction of magnetic transition temperatures is essential, because fixing the lattice to the low- or high-temperature configuration can lead to differences of about 200 K. mHDNNP-driven MD simulations including MC spin flips reveal a small volume increase and the disappearance of the rhombohedral distortion at the Néel temperature of MnO, as the method is able to provide mean geometric and thermodynamic data of the PM phase.

Defects

Magnetic properties of bulk materials differ from those of nanoparticles and doped materials. Surfaces, interfaces, defects, and doping can change the magnetic order leading, e.g., to net surface spins or lower transition temperatures^{81–83}. Surface layers with different stoichiometries can lead to different magnetic orders, such as antiferromagnetic MnO nanoparticles with ferrimagnetic Mn_3O_4 shells⁸⁴. To confirm that such systems can, in principle, be described by mHDNNPs, we predict the impact of Mn vacancies using the mHDNNP. Mn vacancies alter the magnetic order and increase the oxidation states of the remaining Mn ions to compensate the oxygen excess ensuring overall charge neutrality. In principle, it is also possible to study the role of surfaces and doping, but our current parameterization is based on Mn_xO bulk data only, with $x = 0.969$ and 1, and thus the present mHDNNP is not applicable to these situations.

From the heat capacities shown in the Supplementary Fig. 7 and the order parameters in Fig. 6, the Néel temperatures can be determined to be $(298 \pm 1) \text{ K}$ for MnO, $(296 \pm 1) \text{ K}$ for $\text{Mn}_{0.999}\text{O}$, $(275 \pm 1) \text{ K}$ for $\text{Mn}_{0.991}\text{O}$, and $(233 \pm 1) \text{ K}$ for $\text{Mn}_{0.969}\text{O}$. The increasing Mn vacancy concentration decreases the Néel

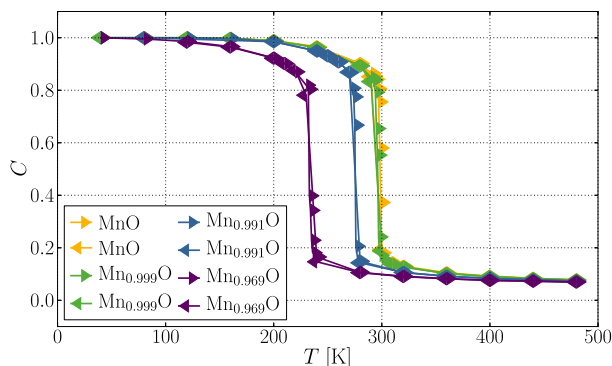


Fig. 6 Impact of Mn vacancies on the temperature dependence of the order parameter. Order parameter C as a function of the temperature T obtained in MDMC simulations for $6 \times 6 \times 6$ supercells of MnO, Mn_{0.999}O, Mn_{0.991}O, and Mn_{0.969}O.

temperature. The more Mn vacancies are present, the lower is the average number of magnetic neighbor ions whose interactions can support the stability of the ordered phase against temperature-induced fluctuations. Consequently, the Néel temperature can be used as a hint for the regularity of a material's crystal structure. This trend is in accordance with experiments that observe an increase of the magnetic transition temperature with higher contents of magnetic ions⁸¹.

Employing the HDNNS, we are able to calculate the absolute magnitude of the atomic spins whose signs are provided in the input of the mHDNNP. In this way, we can investigate the impact of defects on the Mn oxidation states and the magnetization. For this purpose, we optimized the MnO and Mn_{0.969}O structures with AFM-II order using the mHDNNP. For MnO, we observe 16 Mn ions with $M_S = 2.26$ and 16 with $M_S = -2.26$ per magnetic unit cell, whereas for Mn_{0.969}O we find 2 Mn ions with $M_S = 1.83$, 14 with $2.24 \leq M_S \leq 2.26$, and 15 with $-2.26 \leq M_S \leq -2.24$. Consequently, in MnO all Mn ions are Mn^{II} and in Mn_{0.969}O two Mn ions per Mn vacancy are Mn^{III}, whereas the others are Mn^{II}. The difference of the M_S values from the expected values of 2 and $\frac{5}{2}$ is also obtained in the HSE06 reference data. As expected, the excess of O ions due to the Mn vacancy is compensated by increased oxidation of Mn ions. To retain a small net magnetization, Mn^{III} is formed at magnetic lattice sites, which do not have the same spin sign as the missing Mn ion. The sum of the atomic spins per magnetic unit cell is $\bar{M}_S = 1.42$ for Mn_{0.969}O.

MDMC simulations of $6 \times 6 \times 6$ supercells analyzed by the HDNNS reveal that the magnetization of MnO is always zero as shown in Fig. 7, because it is either antiferromagnetic or paramagnetic. However, as the Mn vacancies are placed on an equidistant grid, the constructed Mn_{0.969}O structure is ferrimagnetic at low temperatures. Below the transition temperature, the magnetization is finite but decreases with increasing temperature, because more and more spin flips happen. At the transition temperature of 233 K, a sudden drop to zero occurs as the system becomes paramagnetic. The average number of Mn^{III} ions during all MDMC simulations of Mn_{0.969}O at different temperatures is predicted to be about 1.9 employing the criterion $|M_S| < 2.025$. This value is slightly smaller than the ideal value of 2 such that most of the Mn^{III} are identified correctly. In total, 99.5% of the Mn^{II} and Mn^{III} ions are assigned correctly.

In conclusion, Mn vacancies lead to a reduced Néel temperature and the formation of Mn^{III} ions. The net magnetization induced by Mn vacancies drops to zero at the transition temperature. Consequently, different stoichiometries and associated oxidation states changes are accurately described by the mHDNNP, highlighting the generality and applicability of this method.

Beyond the present work, we expect the mHDNNP method to be a powerful tool for highly accurate, large-scale atomistic

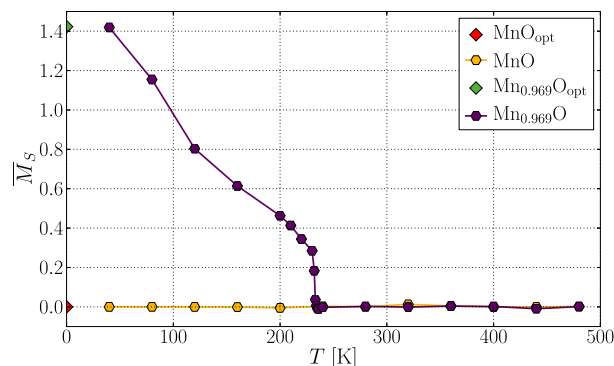


Fig. 7 Temperature dependence of the magnetization. Sum of the atomic spins per magnetic unit cell \bar{M}_S as a function of the temperature T obtained in optimizations and MDMC simulations for $6 \times 6 \times 6$ supercells of MnO and Mn_{0.969}O.

simulations of systems involving different magnetic states, such as a variety of magnetic materials and molecular transition metal complexes containing spin-polarized atoms. Theoretical predictions of the magnetic, geometric, and thermodynamical implications of surfaces, interfaces, defects, and doping can provide interesting control tactics for material properties and finally for technological applications.

METHODS

High-dimensional neural network potential

The MLP used in this work is a second-generation HDNNP⁹. In this method, the potential energy E is constructed as a sum of atomic energy contributions,

$$E = \sum_{m=1}^{N_{\text{elem}}} \sum_{n=1}^{N_{\text{atoms}}^m} E_n^m(\mathbf{G}_n^m), \quad (5)$$

for a system containing N_{elem} elements and N_{atoms}^m atoms of element m . For each element, an individual neural network is trained, which can provide the atomic energy as a function of the local chemical environment and which is evaluated as often as atoms of the respective element are present. The structural descriptors \mathbf{G}_n^m are vectors of many-body ACSFs⁵², which fulfill the mandatory translational, rotational, and permutational invariances of the PES and serve as input vectors of the atomic neural networks. ACSFs describe the local atomic environment as a function of the positions of all neighboring atoms inside a cutoff sphere of radius R_c . To include all energetically relevant interactions, the cutoff radius has to be sufficiently large. Besides the positions of the atoms, only the elements have to be specified leading to a potential being able to describe the making and breaking of bonds. The dimensionality of the ACSF vectors can be predefined for each element individually and does not depend on the atomic environments. This ensures that the number of input neurons of the atomic neural networks remains constant during MD simulations. After optimizing the parameters of the atomic neural networks in a training process using the potential energies and force components of reference systems obtained from DFT, the HDNNP can be applied in large-scale simulations at a small fraction of the computational costs. More details about HDNNPs, their construction, and validation can be found in several reviews^{5,10–12}.

Atom-centered symmetry functions

Two types of ACSFs are most commonly used for the construction of HDNNPs: the radial symmetry functions,

$$G_i^{\text{rad}} = \sum_j e^{-\eta R_{ij}^2} \cdot f_c(R_{ij}), \quad (6)$$

and the angular symmetry functions,

$$G_i^{\text{ang}} = 2^{-\zeta} \sum_j \sum_{k \neq j} [1 + \lambda \cos(\theta_{jkl})]^\zeta \cdot e^{-\eta(R_{ij}^2 + R_{ik}^2 + R_{jk}^2)} \cdot f_c(R_{ij}) \cdot f_c(R_{ik}) \cdot f_c(R_{jk}), \quad (7)$$

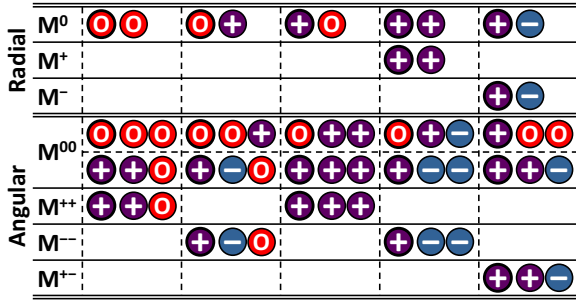


Fig. 8 Interactions resulting in non-zero values of the radial and angular spin-augmentation functions. Red circles with a zero represent atoms with $s = 0$, purple circles with a plus sign atoms with $s = 1$, and blue circles with a minus sign atoms with $s = -1$. The first atom of each entry is the central atom i of the sACSF. The order of the neighbor atoms in the angular interactions is insignificant. The inverse interactions (i.e., switching “+” and “-”) yield the same result and are not shown for clarity.

with the cutoff function

$$f_c(R_{ij}) = \begin{cases} \frac{1}{2} \cos\left(\frac{\pi R_{ij}}{R_c}\right) + \frac{1}{2} & \text{for } R_{ij} \leq R_c \\ 0 & \text{otherwise} \end{cases} \quad (8)$$

R_{ij} is the distance between central atom i and neighboring atom j . θ is the angle $j - i - k$ involving two neighbors j and k . η , λ , and ζ are parameters defining the spatial shapes of the ACSFs. Consequently, the ACSF values only depend on the local geometric environment of the atoms. For multicomponent systems containing several elements, ACSFs for all element combinations are explicitly included. A detailed discussion of the properties of conventional ACSFs and further functional forms can be found in ref. ⁵².

Spin-dependent atom-centered symmetry functions

In many cases, the representation of atomic oxidation and spin states, i.e., the absolute magnitude of the atomic spins, is possible in an indirect way even with conventional ACSFs as shown in our previous studies^{50,69} and in the Supplementary Figs. 1 and 2 of this work. The reason is that different absolute spins in most cases change the geometric structure of the atomic environments. These different environments allow constructing the structure–energy relationship for the ground state by HDNNPs, as long as there is a unique relation between the geometric structure and the electronic configuration. Therefore, ACSFs are able to indirectly capture the energetic effects of the atomic spin magnitude via the geometric structure. Examples are different oxidation states or high- and low-spin states of transition metal ions, for which the different orbital occupations can give rise to structural changes in the local atomic environments such as different bond lengths or Jahn–Teller distortions.

However, the relative parallel and antiparallel arrangement of an atomic spin with respect to the spins of all other magnetic atoms in its environment is not captured by ACSFs. This information is required to distinguish different magnetic configurations and to predict the corresponding potential energies. To extend the applicability to arbitrary collinear spin arrangements, the signs of the atomic spins have to be known explicitly. In principle, also the explicit inclusion of absolute spin values is of interest, but we leave this aspect to future work here.

To describe the magnetic configuration, we introduce an atomic spin coordinate,

$$s_i = \begin{cases} 0 & \text{for } |M_S| < M_S^{\text{thres}} \\ \text{sgn}(M_S) & \text{otherwise} \end{cases}, \quad (9)$$

with

$$M_S = \frac{1}{2}(n_\uparrow - n_\downarrow). \quad (10)$$

M_S is the half-difference of the number of spin-up electrons n_\uparrow and spin-down electrons n_\downarrow of an atom i in a collinear spin-polarized calculation. The atomic spin coordinate is equal to the sign of M_S , unless the absolute atomic spin value is smaller than a threshold value M_S^{thres} . The threshold is introduced to filter out noise in the spin reference data arising from the

ambiguity in assigning spins in electronic structure calculations. In this work, M_S^{thres} is set to 0.25. The set of atomic spin coordinates can represent all possible collinear magnetic configurations enabling to identify ferromagnetic and antiferromagnetic spin arrangements, as well as non-magnetic interactions.

To integrate the spin coordinates into the radial ACSFs, the radial spin-augmentation function (SAF) $M^x(s_i, s_j)$ is employed,

$$G_i^{\text{rad}} = \sum_j M^x(s_i, s_j) \cdot e^{-\eta R_{ij}^2} \cdot f_c(R_{ij}). \quad (11)$$

Different radial SAFs, with $x = 0, +, -$, are used to describe the interactions of same (ferromagnetic interactions) and opposite spin signs (antiferromagnetic interactions), respectively,

$$M^0(s_i, s_j) = 1, \quad (12)$$

$$M^+(s_i, s_j) = \frac{1}{2} |s_i s_j| \cdot |s_i + s_j|, \quad (13)$$

$$M^-(s_i, s_j) = \frac{1}{2} |s_i s_j| \cdot |s_i - s_j|. \quad (14)$$

The radial SAFs $M^x(s_i, s_j)$ are non-zero only for specific combinations of the spin coordinates of the atom pairs as summarized in Fig. 8. This spin augmentation filters the contributions to the radial sACSF in Eq. (11), to distinguish the different magnetic interactions. Only interactions between atoms of parallel spins contribute to sACSFs containing M^+ and only interactions between atoms of antiparallel spins contribute to a sACSFs containing M^- . If $s = 0$ for one or both of the interacting atoms, M^+ and M^- are zero, leaving only a contribution to the sACSF containing M^0 . Taking the absolute values in M^+ and M^- ensures that the descriptor only depends on the relative spin arrangement. In this way, a simultaneous sign change of all atomic spins does not change the value of the sACSFs, ensuring the invariance of the potential energy with respect to the absolute spin arrangement. M^0 is used to describe the non-magnetic, purely geometry-dependent interactions between an $s \neq 0$ atom and an $s = 0$ atom or between two $s = 0$ atoms, which are not included in the other terms.

When constructing an mHDNNP, it is sufficient to use sACSFs only as input for the atomic neural networks of elements exhibiting atoms with non-zero spins. For the atomic neural networks of all other elements, conventional ACSFs can be used, which is equivalent to using only M^0 and M^{00} in the sACSFs of these elements. In the same way as ACSFs, sACSFs are constructed for all element combinations. The choice, which SAFs, i.e., only M^0 or both M^+ and M^- , are required for a given element combination, can be made before constructing the potential, because in most systems the atoms of a given element are either all characterized by $s = 0$ or by $s \neq 0$. For instance, in MnO the Mn atoms exhibit $s \neq 0$ and the O atoms correspond to $s = 0$. Still, the method is also applicable to other systems including partly magnetically active elements. These systems require the use of M^{0^+} (Supplementary Methods) instead of M^0 , to explicitly separate non-magnetic from magnetic interactions, as well as a careful choice of M_S^{thres} to assign physically meaningful spin coordinate values. For element combinations of partly magnetically active elements with (partly) magnetically active elements, all radial SAFs, M^{0^+} , M^+ , and M^- , are required.

In a similar way, angular sACSFs can be defined as

$$G_i^{\text{ang}} = 2^{-\zeta} \sum_j \sum_{k \neq j} M^{\text{ax}}(s_i, s_j, s_k) \cdot [1 + \lambda \cos(\theta_{ijk})]^\zeta \cdot e^{-\eta(R_{ij}^2 + R_{ik}^2 + R_{jk}^2)} \cdot f_c(R_{ij}) \cdot f_c(R_{ik}) \cdot f_c(R_{jk}), \quad (15)$$

containing the angular SAFs $M^{\text{ax}}(s_i, s_j, s_k)$. They allow distinguishing three different interactions of a central $s \neq 0$ atom i with two neighboring $s \neq 0$ atoms j and k : (1) $s_j = s_j = s_k$, (2) $s_j \neq s_j = s_k$, and (3) $s_j = s_j \neq s_k$. The fourth possibility $s_j = s_k \neq s_j$ is equivalent to (3), as the sums over j and k in the angular sACSFs include the interactions $j - i - k$ and $k - i - j$, as both j and k sum over all contributing neighbor atoms to exclude any dependence on the order of the atoms. An efficient separation of these interactions is

given by the functions,

$$M^{00}(s_j, s_j, s_k) = 1, \quad (16)$$

$$M^{++}(s_j, s_j, s_k) = \begin{cases} \frac{1}{2} |s_j| \cdot (|s_j + s_j + s_k| - 1) & \text{for } s_j \neq 0 \wedge s_k \neq 0 \\ \frac{1}{2} |s_j| \cdot |s_j + s_j + s_k| & \text{for } s_j = 0 \wedge s_k = 0, \\ & \text{otherwise} \end{cases}, \quad (17)$$

$$M^{--}(s_j, s_j, s_k) = \begin{cases} \frac{1}{2} |s_j| \cdot (|s_j - s_j - s_k| - 1) & \text{for } s_j \neq 0 \wedge s_k \neq 0 \\ \frac{1}{2} |s_j| \cdot |s_j - s_j - s_k| & \text{for } s_j = 0 \wedge s_k = 0, \\ & \text{otherwise} \end{cases}, \quad (18)$$

$$M^{+-}(s_j, s_j, s_k) = |s_j s_j s_k| \cdot (|s_j + s_j - s_k| - 1), \quad (19)$$

as depicted in Fig. 8. M^{00} is required to describe interactions including more than one $s = 0$ atom or if atom i is $s = 0$. M^{++} , M^{--} , and M^{+-} yield 1 for the interaction types (1), (2), and (3), respectively, and 0 for the other types in case of $s \neq 0$ atoms. For the interactions of two $s \neq 0$ atoms with one $s = 0$ atom, whereby atom i is $s \neq 0$, only M^{++} and M^{--} are required, separating the ferromagnetic and antiferromagnetic interactions similarly as in the radial sACSFs. For systems in which atoms of the same element can be $M_s \neq 0$ and $M_s = 0$, M^{00} (Supplementary Methods) has to be used instead of M^{00} to distinguish non-magnetic and magnetic interactions, and the contributions of M^{++} and M^{--} have to be further split as described in the Supplementary Methods and Supplementary Table 1.

In summary, the combination of both the arrangement of the spins and the geometry-dependent determination of the atomic spin magnitudes is essential to provide all the information necessary to construct reliable PESs in the case of collinear spins. The functional form of the sACSFs might look similar to an Ising-like description at first glance, but the method allows us to study all systems that can also be addressed by collinear DFT calculations, because the mHDNNP simultaneously depends on the geometric structure, indirectly on the atomic spin magnitudes, and on the atomic spin arrangements. For energy and force predictions, the specification of the elements, atomic coordinates, and atomic spin arrangements is required as input covering the full geometric and magnetic configuration space. Both the training of only several selected magnetic configurations and also the training of the full magnetic configuration space including excited magnetic states are possible using mHDNNPs.

High-dimensional neural network spin prediction

HDNNS⁵⁰ can be used to obtain the absolute magnitude of the atomic spins $|M_S|$ as a function of the geometrical structure. From the absolute atomic spins, the oxidation and spin states can be derived. HDNNSs use the same atomic neural network topology as employed in HDNNPs. However, instead of atomic energy contributions, the absolute atomic spins are predicted, which can be directly obtained by the atomic neural networks without taking the sum in Eq. (5).

Simulation analysis

Employing NpT MD simulations, the heat capacity at constant pressure C_p can be obtained from the fluctuations of the total energy,

$$\frac{C_p}{k_B N_A} = \frac{N_{\text{atoms}}}{k_B^2 T^2 N_{\text{steps}}} \sum_{n=0}^{N_{\text{steps}}} [E_{\text{tot}}(n) - \bar{E}_{\text{tot}}]^2, \quad (20)$$

with N_{atoms} atoms in the simulation cell, the mean simulation temperature T , the total energy per atom $E_{\text{tot}}(n)$ as a function of the MD time step n for the total number of simulation steps N_{steps} , and the mean total energy per atom during the simulation \bar{E}_{tot} . k_B is the Boltzmann constant and N_A is the Avogadro constant.

From MC spin-flip simulations, the influence of the magnetic degrees of freedom on the heat capacity at constant volume C_V can be calculated,

$$\frac{C_V}{k_B N_A} = 3 + \frac{N_{\text{atoms}}}{k_B^2 T^2 N_{\text{steps}}} \sum_{n=0}^{N_{\text{steps}}} [E(n) - \bar{E}]^2. \quad (21)$$

The contribution of the atomic motions to C_V is taken into account by $3k_B N_A$ as the atomic motions are not considered in the conventional MC spin-flip simulations. The energy fluctuation is calculated from the potential energies per atom $E(n)$ as a function of the MC step n for the total number of steps N_{steps} compared to the mean potential energy per atom \bar{E} .

The order parameter C is defined as

$$C = \frac{1}{N_{\text{steps}}} \sum_{n=0}^{N_{\text{steps}}} \max \left[\frac{|\mathbf{s}_{(111)} \cdot \mathbf{s}(n)|}{|\mathbf{s}_{(111)}| \cdot |\mathbf{s}(n)|}, \frac{|\mathbf{s}_{(\bar{1}\bar{1}\bar{1})} \cdot \mathbf{s}(n)|}{|\mathbf{s}_{(\bar{1}\bar{1}\bar{1})}| \cdot |\mathbf{s}(n)|}, \frac{|\mathbf{s}_{(1\bar{1}\bar{1})} \cdot \mathbf{s}(n)|}{|\mathbf{s}_{(1\bar{1}\bar{1})}| \cdot |\mathbf{s}(n)|}, \frac{|\mathbf{s}_{(\bar{1}\bar{1}1)} \cdot \mathbf{s}(n)|}{|\mathbf{s}_{(\bar{1}\bar{1}1)}| \cdot |\mathbf{s}(n)|} \right], \quad (22)$$

with \mathbf{s} being the vector of the atomic spin coordinates s_j of each atom i in the simulation cell,

$$\mathbf{s} = (s_1, \dots, s_{N_{\text{atoms}}})^T. \quad (23)$$

The atomic spin coordinates are equal to the sign of the atomic spins as defined in Eq. (9). The vector \mathbf{s} at step n is compared with the vectors $\mathbf{s}_{(111)}$, $\mathbf{s}_{(\bar{1}\bar{1}\bar{1})}$, $\mathbf{s}_{(1\bar{1}\bar{1})}$, and $\mathbf{s}_{(\bar{1}\bar{1}1)}$, which are the possible AFM-II configurations in different spatial orientations. For normalization both vectors are divided by their lengths. The maximum agreements of the relative magnetic configurations at each step are averaged over the simulation length, i.e., the most similar of the four AFM-II spatial orientations is used.

Computational details

The collinear spin-polarized DFT reference calculations were performed employing the Fritz-Haber-Institute ab initio molecular simulations (FHI-aims) code (version 200112.2)^{85,86}. The screened hybrid exchange-correlation functional HSE06 ($\omega = 0.11 a_0$)^{64–66} and the “intermediate” FHI-aims basis set of numeric atom-centered functions excluding the auxiliary 5g hydrogenic functions were employed. A Γ -centered \mathbf{k} -point grid of $2 \times 2 \times 2$ was applied to calculate the $2 \times 2 \times 2$ supercells of MnO (64 atoms without vacancies). The convergence criterion for the self-consistency cycle was set to 10^{-6} eV for the total energies and 10^{-4} eV \AA^{-1} for the forces. Hirshfeld spin moments⁸⁷ were used to determine the atomic spin coordinates. Further details are given in the Supplementary Tables 2 and 3. An extensive benchmark for MnO employing hybrid DFT functionals can be found in our previous work⁸⁸.

The sACSFs were implemented in a modified version of the RuNner code version 1.00^{11,12,89} to construct the mHDNNP and HDNNS. A cutoff radius of $R_c = 10.5 a_0$ was used. A list of the employed parameters of the n_G^m sACSFs for each element m in the mHDNNP is given in Tables 2 and 3. The atomic feed-forward neural networks of mHDNNP and HDNNS consist of n_G^m input neurons, three hidden layers with 20, 15, and 10 neurons, respectively, and one output neuron. The mHDNNP was trained using the cohesive energies, i.e., the total energy minus the sum of the free atom energies, and using atomic force components obtained from DFT calculations of reference structures in different magnetic states. The HDNNS was trained on the absolute values of the Hirshfeld spin moments of the same reference structures. Ninety percent of these data were used for training the neural networks. The remaining data were used as test set. Further details about the training can be found in the Supplementary Tables 4 and 5^{50,69}.

mHDNNP-driven MD simulations in combination with MC spin flips were carried out using the Large-scale Atomic/Molecular Massively Parallel Simulator^{90,91} and the neural network potential package (n2p2)⁹² as library for potentials generated with RuNner. The n2p2 code was modified to enable the usage of sACSFs. The MD simulations with MC spin flips (MDMC) employed $6 \times 6 \times 6$ Mn_xO supercells referring to the geometric unit cell, with $x = 0.969, 0.991, 0.999, 1$, i.e., 1701, 1720, 1727, and 1728 atoms. They were run in the NpT ensemble at a pressure of $p = 1$ bar with a time step of 1 fs applying the Nosé–Hoover thermostat and barostat with coupling constants of 0.1 and 1 ps, respectively^{93,94}. MC spin flips were performed after each time step. The spin-flip rate is not set to measured or calculated rates to study dynamic properties but to sample the thermodynamic equilibrium efficiently. In all MDMC simulations, the

Table 2. Employed radial sACSFs.

$i-j$	M^x	$\eta [a_0^{-2}]$
O-O	M^0	0, 0.00117, 0.00246, 0.00389, 0.00550
O-Mn	M^0	0, 0.00369, 0.00882, 0.01636, 0.02809
Mn-O	M^0	0, 0.00369, 0.00882, 0.01636, 0.02809
Mn-Mn	M^+, M^-	0, 0.00085, 0.00176, 0.00274, 0.00381

All combinations of SAFs and symmetry function parameters are applied for the given element pairs using $R_c = 10.5 a_0$.

Table 3. Employed angular sACSFs.

$i-j-k$	$M^{\alpha\alpha}$	λ	ζ
O-O-O	M^{00}	-1, 1	1, 2, 4, 16
O-O-Mn	M^{00}	-1, 1	1, 2, 4, 16
O-Mn-Mn	M^{00}	-1, 1	1, 2, 4, 16
Mn-O-O	M^{00}	-1, 1	1, 2, 4, 16
Mn-O-Mn	M^{++}, M^-	-1, 1	1, 2, 4, 16
Mn-Mn-Mn	M^{++}, M^-, M^{+-}	-1, 1	1, 2, 4, 16

All combinations of SAFs and symmetry function parameters are applied for the given element combinations using $\eta = 0$ a_0^{-2} and $R_c = 10.5$ a_0 .

system was equilibrated for 1 ns before the acquisition period of 10 ns. In all conventional MC spin-flip simulations, i.e., no MD steps in between the MC spin flips, the equilibration was performed for 10^6 steps and the acquisition consisted of 10^7 steps.

DATA AVAILABILITY

The data sets generated and analyzed during the current study are available from the corresponding author on reasonable request.

CODE AVAILABILITY

The modified versions of RuNNer and n2p2 to enable the usage of sACSFs are available from the corresponding author on reasonable request. The modifications will be implemented in coming release versions under the GPL3 license.

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AUTHOR CONTRIBUTIONS

M.E. conceived the sACSF approach and initiated the project. M.E. worked out and implemented the practical algorithms and performed all calculations. M.E. and J.B. contributed ideas to the project and analyzed the results. M.E. wrote the initial version of the manuscript and prepared the figures. M.E. and J.B. jointly edited the manuscript.

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